

# ローレンツ対称性を保つタイプ**IIB**行列模型の 新しい定義とその必要性

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Ref.) Asano, JN, Piensuk, Yamamori, arXiv:2404. 14045

Asano, JN, Piensuk, Yamamori, in preparation

Chou, JN, Tripathi, in preparation

# 0. Introduction

# type IIB (or IKKT) matrix model

Ishibashi-Kawai-Kitazawa-Tsuchiya,

Nucl.Phys.B 498 (1997) 467, hep-th/9612115 [hep-th]

- a **nonperturbative** formulation of **superstring** theory  
“**lattice gauge theory**” of **everything** (matter, force and space-time)

$$S_b = -\frac{1}{4}N \operatorname{tr}([A_\mu, A_\nu][A^\mu, A^\nu])$$
$$S_f = -\frac{1}{2}N \operatorname{tr}(\Psi_\alpha (C \Gamma^\mu)_{\alpha\beta} [A_\mu, \Psi_\beta])$$

( 0-dimensional reduction  
of 4D  $\mathcal{N} = 4$  SYM )

$N \times N$  Hermitian matrices      **SO(9,1) Lorentz symmetry**

$A_\mu$  ( $\mu = 0, \dots, 9$ )      Lorentz vector  
 $\Psi_\alpha$  ( $\alpha = 1, \dots, 16$ )      Majorana-Weyl spinor

Lorentzian metric  $\eta = \operatorname{diag}(-1, 1, \dots, 1)$   
is used to raise and lower indices.

- Unlike AdS/CFT, not only **space** but also **time emerge**  
as the **eigenvalue distribution** of the 10 bosonic matrices.

maximal SUSY (incl. translation :  $A_\mu \mapsto A_\mu + \alpha_\mu \mathbf{1}$  )

# classical solutions

- Eq. of motion :  $[A^\nu, [A_\nu, A_\mu]] = 0$

Classical solutions are exhausted  
by the diagonal ones ( $[A_\mu, A_\nu] = 0$ )

See Appendix A of  
H. C. Steinacker, JHEP 02, 033,  
arXiv:1709.10480 [hep-th].



Add a Lorentz invariant “mass” term to the IKKT action.

$$S_m = -\frac{1}{2} N \gamma \operatorname{tr}(A_\mu A^\mu) = \frac{1}{2} N \gamma \left\{ \operatorname{tr}(A_0)^2 - \operatorname{tr}(A_i)^2 \right\}$$

$$S_b = \frac{1}{4} N \operatorname{tr}(F_{\mu\nu} F^{\mu\nu}) = \frac{1}{4} N \left\{ -2 \operatorname{tr}(F_{01})^2 + \operatorname{tr}(F_{ij})^2 \right\}$$

$$F_{\mu\nu} = i [A_\mu, A_\nu] \\ \text{(Hermitian)}$$

$$\text{Eq. of motion : } [A^\nu, [A_\nu, A_\mu]] - \gamma A_\mu = 0$$

Many classical solutions representing  
expanding space-time appear.

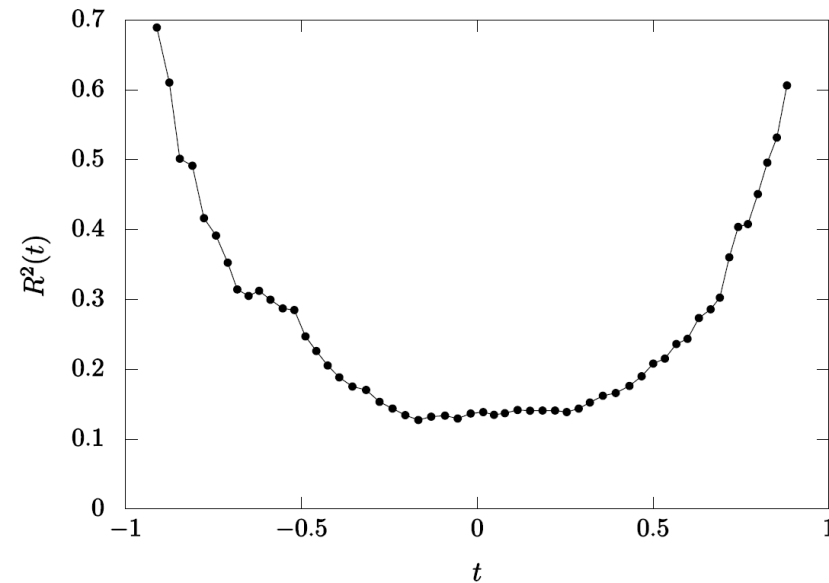
Kim-J.N.-Tsuchiya, 1208.0711  
Sperling-Steinacker 1901.03522

# typical classical solutions

Eq. of motion :  $[A^\nu, [A_\nu, A_\mu]] - \gamma A_\mu = 0$

Hatakeyama-Matsumoto-J.N.-  
Tsuchiya-Yosprakob,  
*PTEP* 2020 (2020) 4, 043B10

- $A_\mu = 0$  is always a solution.  
(trivial solution)
- Typical Hermitian  $A_\mu$  solutions show expanding behavior for  $\gamma > 0$  but **not** for  $\gamma < 0$  !
- However, space-time dimensionality is not determined **at the classical level**.



We have to investigate the partition function including the effects of fermions in the 1)  $N \rightarrow \infty$ , 2)  $\gamma \rightarrow +0$  lim. to see if (3+1)D expanding space-time appears.

- $\gamma = 0$  is a “strong coupling” limit.

$$Z = \int dA e^{i(A^4 + \gamma A^2)} \quad A_\mu = \sqrt{|\gamma|} \tilde{A}_\mu$$

$$= \int dA e^{i\gamma^2(\tilde{A}^4 + \tilde{A}^2)} \quad \gamma^2 \Leftrightarrow \frac{1}{\hbar}$$

Quantum effects become important.  
The role of SUSY.

# partition function of the type IIB matrix model

$$\begin{aligned} Z &= \int dA d\Psi e^{i(S_b + S_m + S_f)} \\ &= \int dA \underbrace{e^{i(S_b + S_m)}}_{\text{pure phase factor}} \underbrace{\text{Pf} \mathcal{M}(A)}_{\text{polynomial in } A} \end{aligned}$$

The partition function is NOT absolutely convergent.

- As a regularization, it was proposed to add **convergence factors**.

$$Z = \int dA e^{i(S_b + S_m)} \text{Pf} \mathcal{M}(A)$$

Anagnostopoulos-Azuma-Hatakeyama-  
Hirasawa-J.N.-Papadoudis-Tsuchiya,  
in preparation

$$S_m^{(\varepsilon)} = \frac{1}{2} N \gamma \left\{ e^{i\varepsilon} \text{tr}(A_0)^2 - e^{-i\varepsilon} \text{tr}(A_i)^2 \right\} \quad \text{This breaks Lorentz symmetry!}$$

In fact, the partition function diverges in the  $\varepsilon \rightarrow 0$  limit due to noncompact flat directions, and the cutoff artifact remains.

# What we do in this talk

- We study N=2 bosonic model with the cutoff **nonperturbatively** by **1/D expansion** and **the generalized thimble method (GTM)**.
- In particular, **“classicalization”** occurs in the  $\epsilon \rightarrow 0$  limit due to **an artifact of the Lorentz symmetry breaking cutoff**.



- We propose a new definition of the type IIB matrix model that **respects Lorentz symmetry using Faddeev-Popov gauge fixing**.
- Our results for the gauge-fixed model (**obtained by GTM**) show very different behaviors.

⌈ Once we understand the N=2 bosonic model completely, we just have to do the same things for larger N with SUSY. ⌋

# Plan of the talk

0. Introduction
1. The need for “gauge-fixing” in the toy model
2. Classical solutions in  $N=2$  bosonic model
3.  $1/D$  expansion and MC studies of the cutoff model
4. MC studies of the “gauge-fixed” model
5. Summary and discussions



# 1. The need for gauge fixing in the toy model

Asano, JN, Piensuk, Yamamori, arXiv:2404. 14045

# a toy model with Lorentz symmetry


type IIB matrix model

$$A_\mu = \sum_{a=1}^{N^2-1} A_\mu^a t^a \quad \text{basis of traceless Hermitian matrices}$$

a model of  $(N^2 - 1)$  Lorentz vectors

a toy model with a single Lorentz vector  $x_\mu$  ( $\mu = 0, 1, \dots, d$ )

$$Z = \int dx e^{-\frac{1}{2}\gamma(x_\mu x^\mu + 1)^2} \quad \gamma > 0$$

Saddle points : (i)  $(x_0)^2 - (x_i)^2 = 1$   Saddle points of this type are related with each other by Lorentz transformation.  
(ii)  $x_\mu = 0$

**Z diverges due to flat directions**

● cutoff model

$$Z_\varepsilon = \int dx e^{-\frac{1}{2}\gamma(x_\mu x^\mu + 1)^2 - \varepsilon(x_0)^2 - \varepsilon(x_i)^2} \quad \lim_{\varepsilon \rightarrow 0} Z_\varepsilon = \infty$$

# classicalization in the cutoff model

- cutoff model

$$Z_\epsilon = \int dx e^{-\frac{1}{2}\gamma(x_\mu x^\mu + 1)^2 - \epsilon(x_0)^2 - \epsilon(x_i)^2}$$

- One can solve this model by introducing an auxiliary variable k

$$Z_\epsilon = \frac{1}{\sqrt{2\pi\gamma}} \int dk dx e^{-\frac{1}{2\gamma}k^2 + ik(x_\mu x^\mu + 1) - \epsilon(x_0)^2 - \epsilon(x_i)^2}$$

$$Z_\epsilon = \frac{1}{\sqrt{2\pi\gamma}} \int dk e^{-\frac{1}{2\gamma}k^2 + ik} \sqrt{\frac{\pi}{ik + \epsilon}} \left( \sqrt{\frac{\pi}{-ik + \epsilon}} \right)^d$$

$$= \mathcal{N} \int dk e^{-S_{\text{eff}}(k)} \quad S_{\text{eff}}(k) = \frac{1}{2\gamma}k^2 - ik + \frac{1}{2} \log(ik + \epsilon) + \frac{d}{2} \log(-ik + \epsilon)$$

$$0 = \frac{dS_{\text{eff}}(k)}{dk} = \frac{1}{\gamma}k - i + \frac{i}{2} \frac{1}{ik + \epsilon} - \frac{id}{2} \frac{1}{-ik + \epsilon}$$

dominant saddle point

$$k^{(0)} \simeq i \frac{d-1}{d+1} \epsilon + i \frac{8d}{(d+1)^3} \epsilon^2 + \mathcal{O}(\epsilon^3)$$

$$Z_\epsilon \sim \epsilon^{-\frac{d+1}{2}} (\times \epsilon) \sim \epsilon^{-\frac{d-1}{2}}$$

Partition function diverges for  $\epsilon \rightarrow 0$

$$\langle k \rangle_\epsilon = i\gamma \langle (x_\mu x^\mu + 1) \rangle_\epsilon$$

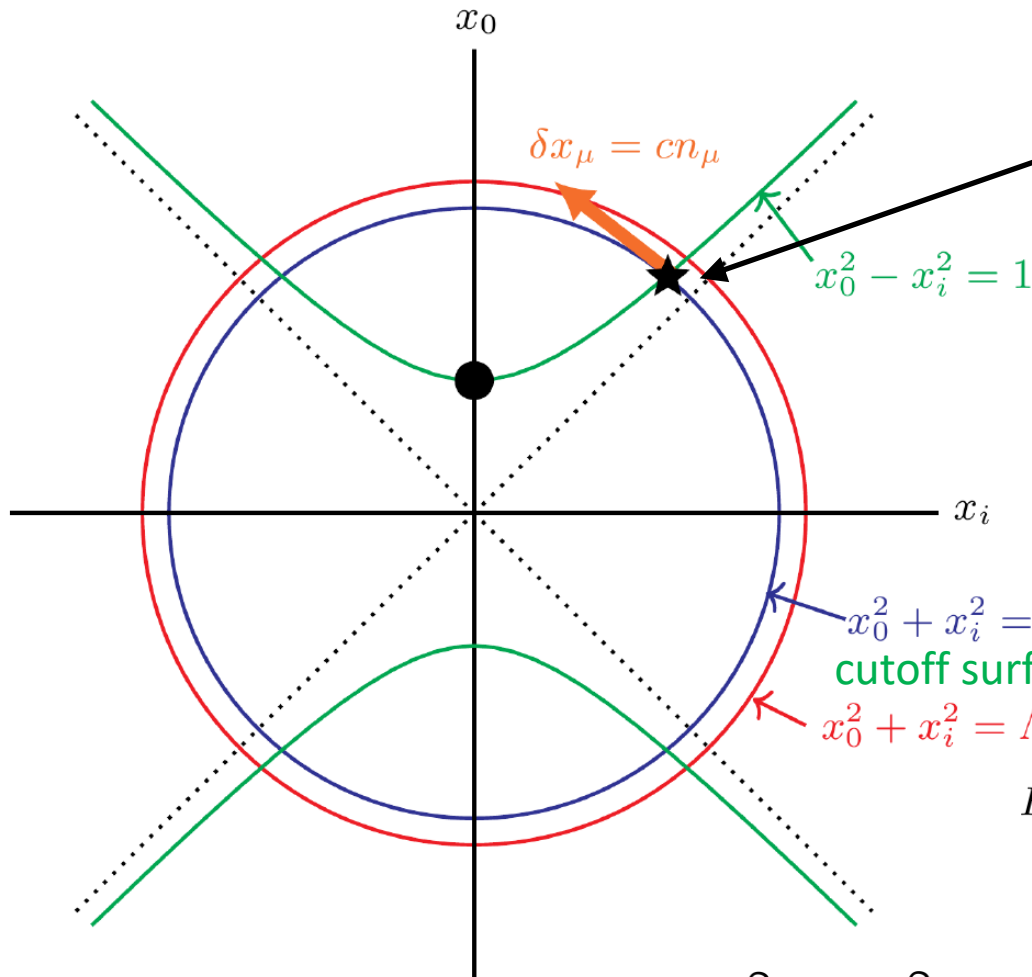
$\epsilon \rightarrow 0$

0

$$- \lim_{\epsilon \rightarrow 0} \langle x_\mu x^\mu \rangle_\epsilon = 1$$

classicalization

# classicalization : an artifact of the cutoff



(Lorentz boosted) saddle point  
 $(x_0, x_1) = (\cosh \sigma, \sinh \sigma)$

$$(\cosh \sigma)\delta x_0 + (\sinh \sigma)\delta x_1 = 0$$

$$\begin{pmatrix} \delta x_0 \\ \delta x_1 \end{pmatrix} = \frac{c}{\sqrt{\cosh 2\sigma}} \begin{pmatrix} \sinh \sigma \\ -\cosh \sigma \end{pmatrix}$$

magnitude of the fluctuations

$$H_{\mu\nu} = \frac{\partial^2 S(x)}{\partial x_\mu \partial x_\nu} = \gamma x^\mu x^\nu$$


$$= \gamma \begin{pmatrix} \cosh^2 \sigma & -\cosh \sigma \sinh \sigma \\ -\cosh \sigma \sinh \sigma & \sinh^2 \sigma \end{pmatrix}$$

$$\delta S = \delta x_\mu H_{\mu\nu}(x) \delta x_\nu = \frac{c^2 \gamma \sinh^2(2\sigma)}{\cosh(2\sigma)}$$

➔  $|c| \lesssim e^{-\sigma} \sqrt{\frac{2}{\gamma}}$  at large  $\sigma$

Classicalization occurs  
 due to the cutoff artifact!

# “gauge-fixing” the Lorenz symmetry

- “Gauge fixing” condition : minimize:  $(x_0)^2$  w.r.t. Lorenz tr.
- 
- $$\begin{pmatrix} x'_0 \\ x'_j \end{pmatrix} = \begin{pmatrix} \cosh \sigma & \sinh \sigma \\ \sinh \sigma & \cosh \sigma \end{pmatrix} \begin{pmatrix} x_0 \\ x_j \end{pmatrix}$$
- $x_0 x_j = 0$  for all  $j$

$$Z = \int dx e^{-\frac{1}{2}\gamma(x_\mu x^\mu + 1)^2} \Delta_{\text{FP}}[x] \prod_{j=1}^d \delta(x_0 x_j)$$

$$\Delta_{\text{FP}}[x] = \det \Omega, \quad \Omega_{ij} = (x_0)^2 \delta_{ij} + x_i x_j$$

- In fact, the time-like region dominates over the space-like region.
- $(x_j = 0)$

$(x_0 = 0)$

$$Z_{\text{g.f.}} = \int dx_0 |x_0|^d e^{-\frac{1}{2}\gamma\{-(x_0)^2 + 1\}^2}$$


$$-\langle x_\mu x^\mu \rangle = 1 + \frac{d-1}{2\gamma} + \dots$$

Fluctuations exist for finite  $\gamma$

The classicalization in the cutoff model is an artifact of Lorenz symmetry breaking that remains in the  $\varepsilon \rightarrow 0$  limit.

# a new definition of type IIB matrix model

- “Gauge fixing” condition : minimize:  $\text{tr}(A_0)^2$  w.r.t. Lorentz tr.



$$\begin{pmatrix} A'_0 \\ A'_j \end{pmatrix} = \begin{pmatrix} \cosh \sigma & \sinh \sigma \\ \sinh \sigma & \cosh \sigma \end{pmatrix} \begin{pmatrix} A_0 \\ A_j \end{pmatrix}$$

$\text{tr}(A_0 A_j) = 0$  for all  $j$

$$Z = \int dA e^{i(S_b + S_m)} \Delta_{\text{FP}}[A] \prod_{j=1}^d \delta(\text{tr}(A_0 A_j))$$
$$\Delta_{\text{FP}}[A] = \det \Omega, \quad \Omega_{ij} = \text{tr}(A_0)^2 \delta_{ij} + \text{tr}(A_i A_j)$$

- We still have to take care of **the oscillating integral** by introducing the convergence factor:

$$S_m^{(\varepsilon)} = \frac{1}{2} N \gamma \left\{ e^{i\varepsilon} \text{tr}(A_0)^2 - e^{-i\varepsilon} \text{tr}(A_i)^2 \right\}$$

Unlike the gauge-unfixed model,  
the partition function is finite in the  $\varepsilon \rightarrow 0$  limit.

## 2. Classical solutions in $N=2$ bosonic model

Asano, JN, Piensuk, Yamamori, in preparation

# classical solutions for the N=2 case

classical EOM :  $[A^\nu, [A_\nu, A_\mu]] - \gamma A_\mu = 0$

For N=2, we can obtain all the real solutions up to  $SO(9,1) \times SU(2)$  sym.

$\gamma < 0$	$A_\mu = 0$		
$\gamma > 0$	$A_\mu = 0$	$A_\mu = \begin{cases} \sqrt{\frac{\gamma}{8}} \sigma_\mu & \mu = 1, 2, 3 \\ 0 & \text{otherwise} \end{cases}$	$A_\mu = \begin{cases} \sqrt{\frac{\gamma}{4}} \sigma_\mu & \mu = 1, 2 \\ 0 & \text{otherwise} \end{cases}$
	(trivial solution)	(Pauli solution)	(squashed Pauli solution)

remaining symmetries

$SO(9,1) \times SU(2)$   
unbroken

diagonal subgroup of  
 $SO(3) \times SU(2)$

diagonal subgroup of  
 $SO(2) \times U(1)$

**Nontrivial real solutions exist only for  $\gamma > 0$  .**



# comments on complex solutions

- real solutions are exhausted (up to symmetries) by

$\gamma < 0$	$A_\mu = 0$
$\gamma > 0$	$A_\mu = 0$ $A_\mu = \begin{cases} \sqrt{\frac{\gamma}{8}} \sigma_\mu & \mu = 1, 2, 3 \\ 0 & \text{otherwise} \end{cases}$ $A_\mu = \begin{cases} \sqrt{\frac{\gamma}{4}} \sigma_\mu & \mu = 1, 2 \\ 0 & \text{otherwise} \end{cases}$

(trivial solution)

(Pauli solution)

(squashed Pauli solution)

- In fact, there are many **complex solutions**.

e.g.)

$\gamma < 0$	$A_\mu = \begin{cases} i\sqrt{\frac{ \gamma }{8}} \sigma_\mu & \mu = 1, 2, 3 \\ 0 & \text{otherwise} \end{cases}$	$A_\mu = \begin{cases} i\sqrt{\frac{ \gamma }{4}} \sigma_\mu & \mu = 1, 2 \\ 0 & \text{otherwise} \end{cases}$
$\gamma > 0$	$\begin{cases} A_0 = i\sqrt{\frac{\gamma}{8}} \sigma_1 \\ A_1 = \sqrt{\frac{\gamma}{8}} \sigma_2 \\ A_2 = \sqrt{\frac{\gamma}{8}} \sigma_3 \\ A_i = 0 \text{ for } i \geq 3 \end{cases}$	$\begin{cases} A_0 = i\sqrt{\frac{\gamma}{4}} \sigma_1 \\ A_1 = \sqrt{\frac{\gamma}{4}} \sigma_2 \\ A_i = 0 \text{ for } i \geq 2 \end{cases}$

These are all **irrelevant** from the viewpoint of the Picard-Lefschetz theory.



thimble

# Picard-Lefschetz theory

(multi-dimensional version of the steepest descent method)

$$Z = \int_{\mathbb{R}^N} dx e^{-S(x)} \quad S(x) \in \mathbb{C} \quad (\text{oscillating integral})$$

relevant saddles

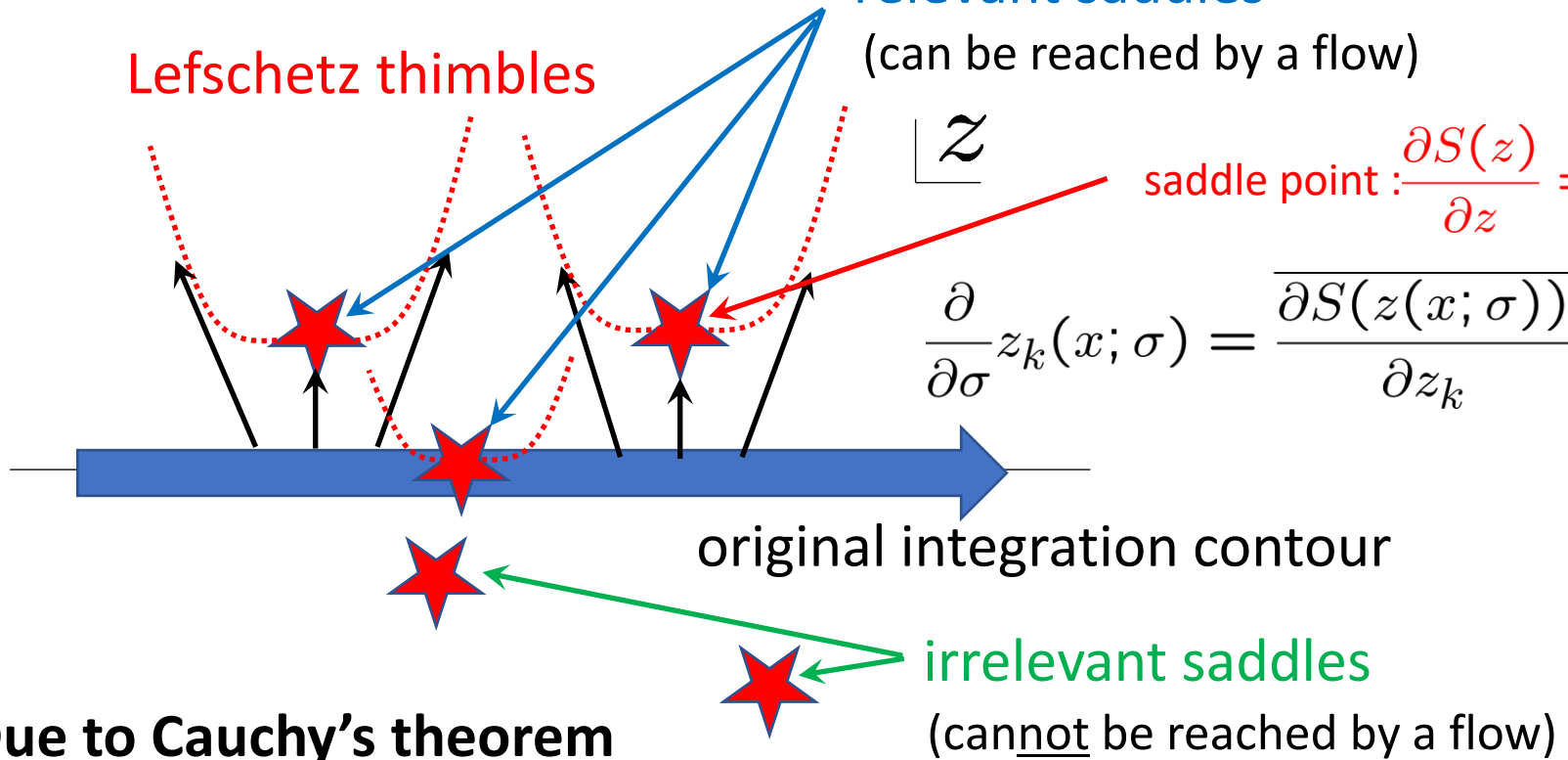
(can be reached by a flow)

Lefschetz thimbles

$\mathcal{Z}$

saddle point:  $\frac{\partial S(z)}{\partial z} = 0$

$$\frac{\partial}{\partial \sigma} z_k(x; \sigma) = \frac{\partial S(z(x; \sigma))}{\partial z_k}$$



original integration contour

irrelevant saddles

(cannot be reached by a flow)

**Due to Cauchy's theorem**

An oscillating integral can be evaluated by summing over all the thimbles associated with the relevant saddle points.

an important property of the flow

Note that  $S(x)$  is complex!

$$\begin{aligned} \frac{d}{d\sigma} S(z(x; \sigma)) &= \frac{\partial S(z(x; \sigma))}{\partial z_k} \frac{\partial z_k(x; \sigma)}{\partial \sigma} \\ &= \frac{\partial S(z(x; \sigma))}{\partial z_k} \overline{\frac{\partial S(z(x; \sigma))}{\partial z_k}} \\ &= \left| \frac{\partial S(z(x; \sigma))}{\partial z_k} \right|^2 \end{aligned}$$

real positive !



Real part of the action increases along the flow, while the imaginary part is kept constant.

complex saddles are irrelevant in the bosonic model

$$Z = \int dA e^{-S[A]} \quad S[A] = -i(A^4 + \gamma A^2)$$

$\text{Re} S[A] = 0$  for real configurations

$\text{Re} S[A] > 0$  required for complex saddles to be relevant

● complex solutions

$\text{Re} S[A] = 0$

e.g.)  $\gamma < 0$

$$A_\mu = \begin{cases} i\sqrt{\frac{|\gamma|}{8}} \sigma_\mu & \mu = 1, 2, 3 \\ 0 & \text{otherwise} \end{cases} \quad A_\mu = \begin{cases} i\sqrt{\frac{|\gamma|}{4}} \sigma_\mu & \mu = 1, 2 \\ 0 & \text{otherwise} \end{cases}$$

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$$\gamma > 0 \quad \begin{cases} A_0 = i\sqrt{\frac{\gamma}{8}} \sigma_1 \\ A_1 = \sqrt{\frac{\gamma}{8}} \sigma_2 \\ A_2 = \sqrt{\frac{\gamma}{8}} \sigma_3 \\ A_i = 0 \quad \text{for } i \geq 3 \end{cases} \quad \begin{cases} A_0 = i\sqrt{\frac{\gamma}{4}} \sigma_1 \\ A_1 = \sqrt{\frac{\gamma}{4}} \sigma_2 \\ A_i = 0 \quad \text{for } i \geq 2 \end{cases}$$

These are all **irrelevant** from the viewpoint of the Picard-Lefschetz theory.

### 3. $1/D$ expansion and MC studies of the cutoff model

Asano, JN, Piensuk, Yamamori, in preparation

# 1/D expansion

Used in the Euclidean model  
without the mass term

Hotta-J.N.-Tsuchiya ('98)

$$A_\mu = \sum_{a=1}^{N^2-1} A_\mu^a t^a \quad h_{ab} \sim A_\mu^a A^{\mu b}$$

$$\begin{aligned} Z &= \int dA e^{i(A^4 + \gamma A^2)} \\ &= \int dh \int dA e^{i(h^2 + hA^2 + \gamma A^2)} \\ &= \int dh e^{ih^2 - \frac{D}{2} \log \det K} \\ &= \int d\tilde{h} e^{-D S_{\text{eff}}[\tilde{h}]} \end{aligned}$$

For the moment, we omit  
the convergence factors.

$$\begin{aligned} \tilde{h}_{ab} &= \frac{1}{\sqrt{D}} h_{ab} \\ \tilde{\gamma} &= \frac{1}{\sqrt{D}} \gamma \end{aligned}$$

$D$  appears here only as a parameter.

At large  $D$  with fixed  $\tilde{\gamma}$ ,

$$\frac{\partial S_{\text{eff}}[\tilde{h}]}{\partial \tilde{h}_\mu} = 0 \quad \longrightarrow \quad \tilde{h} + iK[\tilde{h}]^{-1} = 0$$

# Large D saddles for N=2 bosonic model

Large D SPE :  $\tilde{h} + iK[\tilde{h}]^{-1} = 0$

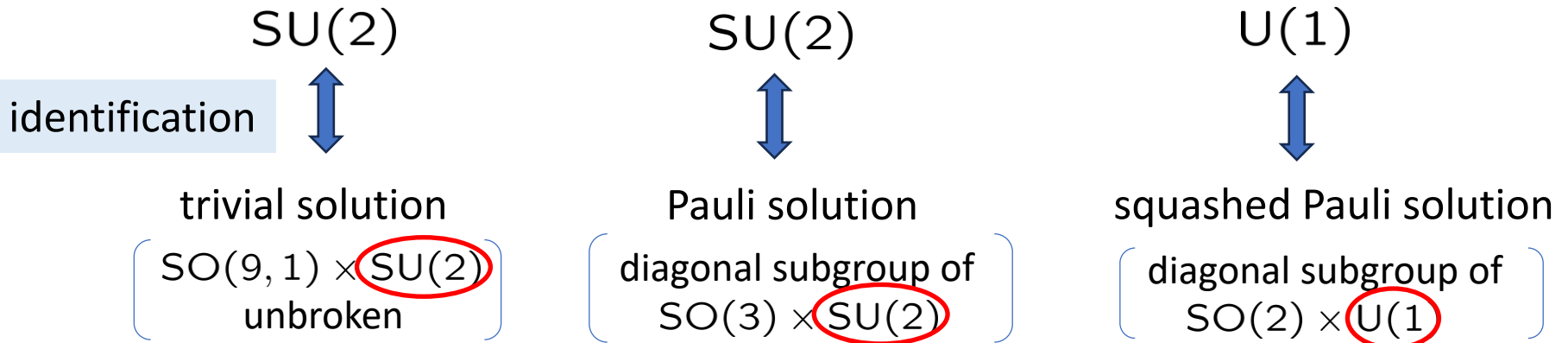
For N=2, we can obtain all the **relevant** saddle points up to symmetries.

$\gamma < 0$	$\tilde{h} = v^{(-)} \mathbf{1}$		
$\gamma > 0$	$\tilde{h} = v^{(-)} \mathbf{1}$	$\tilde{h} = v^{(+)} \mathbf{1}$	$\tilde{h} = \tilde{\gamma} \text{diag} \left( 1, 1, \frac{i}{\tilde{\gamma}^2} \right)$

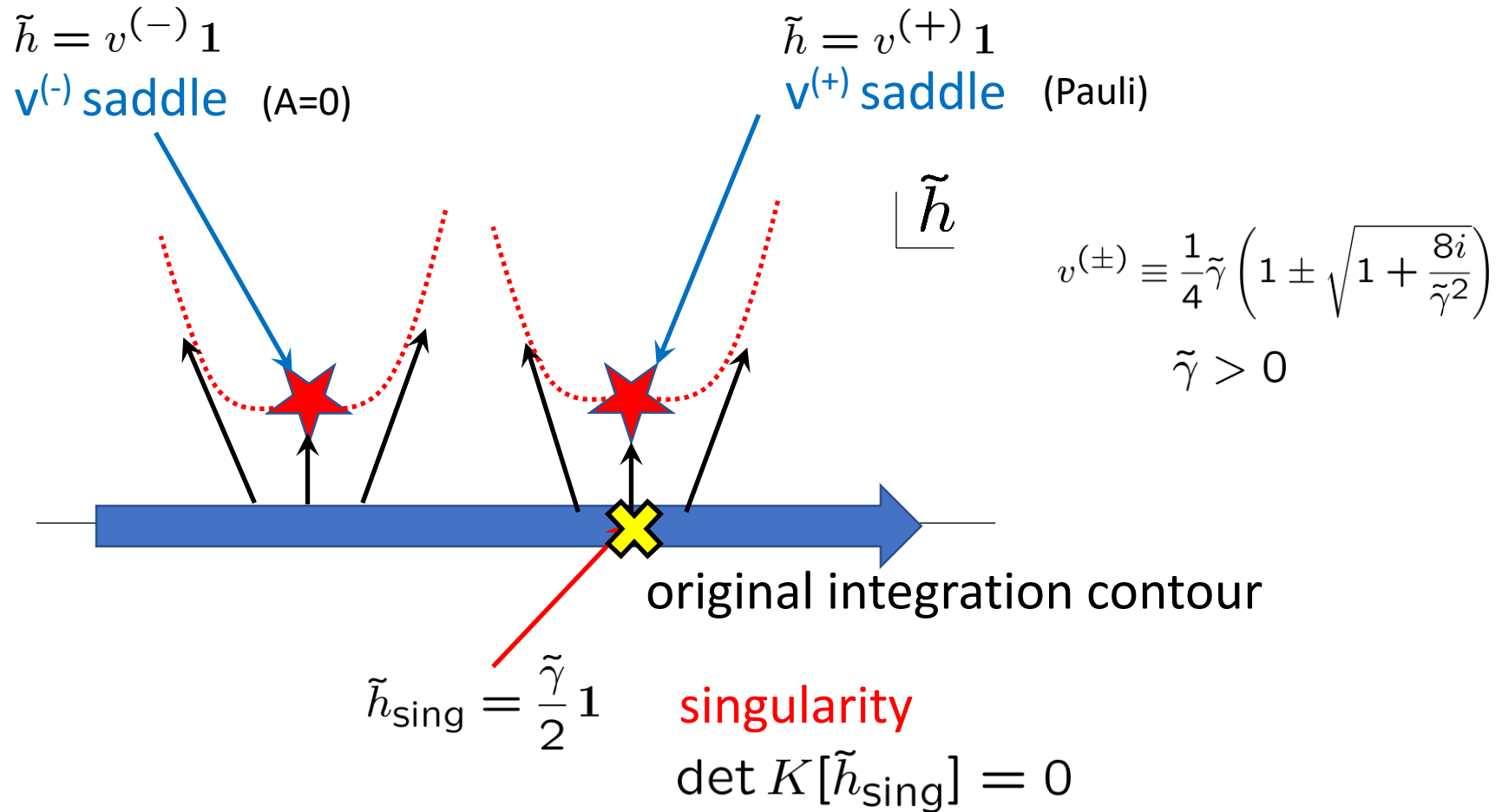
$$v^{(\pm)} \equiv \frac{1}{4} \tilde{\gamma} \left( 1 \pm \sqrt{1 + \frac{8i}{\tilde{\gamma}^2}} \right)$$

These are complex saddles, but not necessarily irrelevant.

remaining symmetries



# Singularity on the real axis



This simply reflects the fact that the partition function is not well defined as it is.

Also true for the SO(D) invariant case !



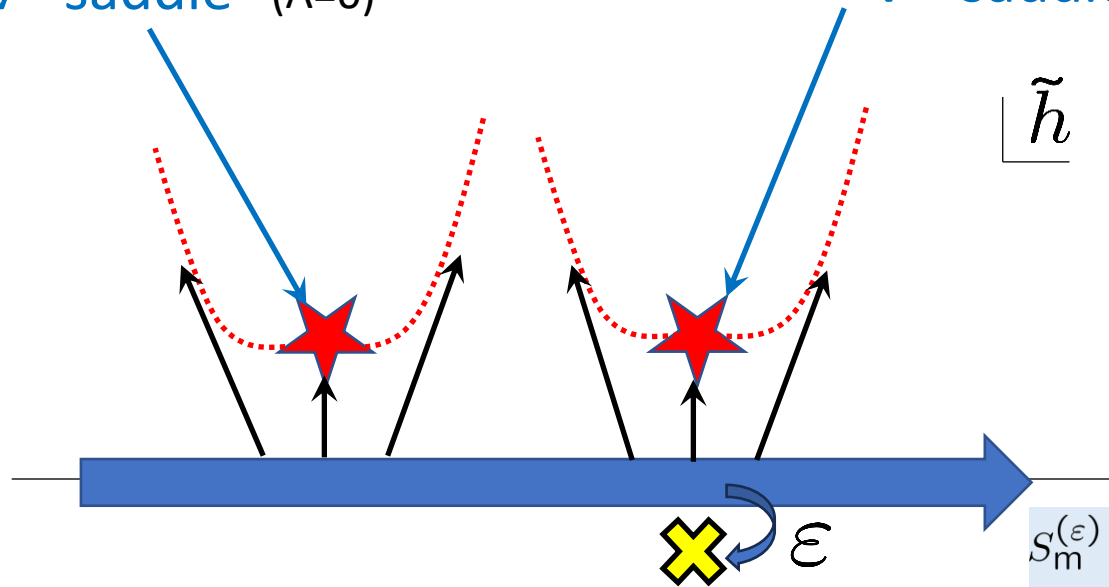
# The case of SO(D) symmetric model obtained by replacing $A_0 = iA_D$

$$Z^{(+)} \sim e^{-\frac{3}{8}iD\tilde{\gamma}^2} \left(\frac{\tilde{\gamma}}{2}\right)^{\frac{3}{2}D}$$

(large D)

$\tilde{h} = v^{(-)} \mathbf{1}$   
 $v^{(-)}$  saddle (A=0)

$\tilde{h} = v^{(+)} \mathbf{1}$   
 $v^{(+)}$  saddle (Pauli)



$\tilde{h}$

$$v^{(\pm)} \equiv \frac{1}{4}\tilde{\gamma} \left(1 \pm \sqrt{1 + \frac{8i}{\tilde{\gamma}^2}}\right)$$

$$\tilde{\gamma} > 0$$

$$S_m^{(\varepsilon)} = -\frac{1}{2} N\gamma e^{-i\varepsilon} \left\{ \text{tr}(A_D)^2 + \text{tr}(A_i)^2 \right\}$$

The  $v^{(+)}$  saddle becomes **relevant** and the associated partition function becomes finite in the  $\varepsilon \rightarrow 0$  limit.

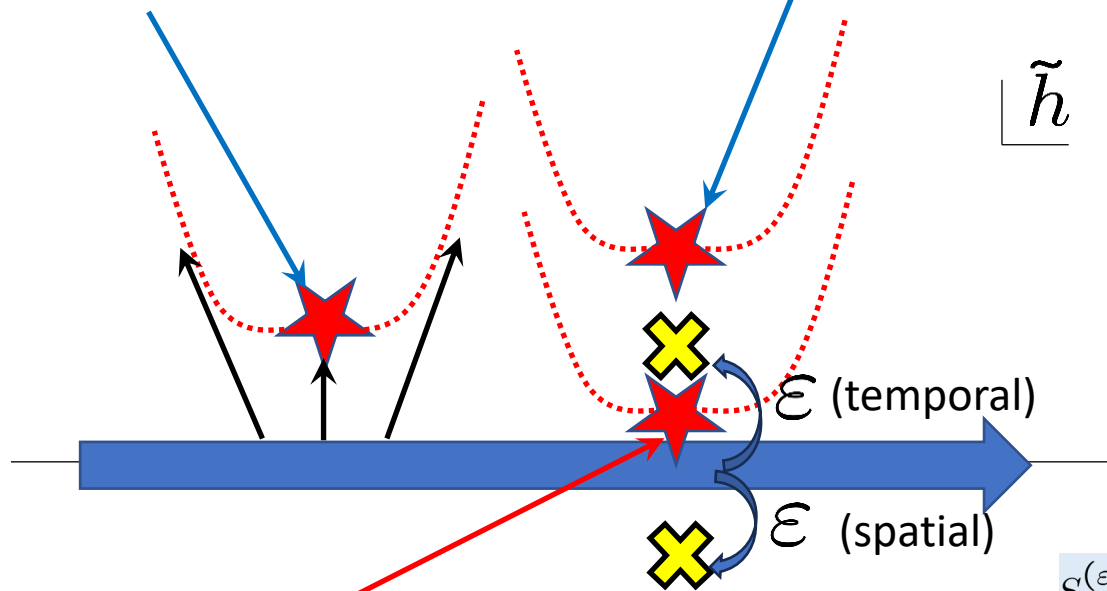
For  $\tilde{\gamma} < 0$ , the  $v^{(+)}$  saddle becomes **irrelevant** since the singularity is shifted in the opposite direction.

consistent with the existence of the Pauli solution only for  $\tilde{\gamma} > 0$ .

# The case of Lorentz symmetric model

$\tilde{h} = v^{(-)} \mathbf{1}$   
 $v^{(-)}$  saddle (A=0)

$\tilde{h} = v^{(+)} \mathbf{1}$   
 $v^{(+)}$  saddle becomes irrelevant!



$$v^{(\pm)} \equiv \frac{1}{4} \tilde{\gamma} \left( 1 \pm \sqrt{1 + \frac{8i}{\tilde{\gamma}^2}} \right)$$

$$\tilde{\gamma} > 0$$

Convergence factor acts on space and time differently.

$$S_m^{(\varepsilon)} = \frac{1}{2} N \gamma \left\{ e^{i\varepsilon} \text{tr}(A_0)^2 - e^{-i\varepsilon} \text{tr}(A_i)^2 \right\}$$

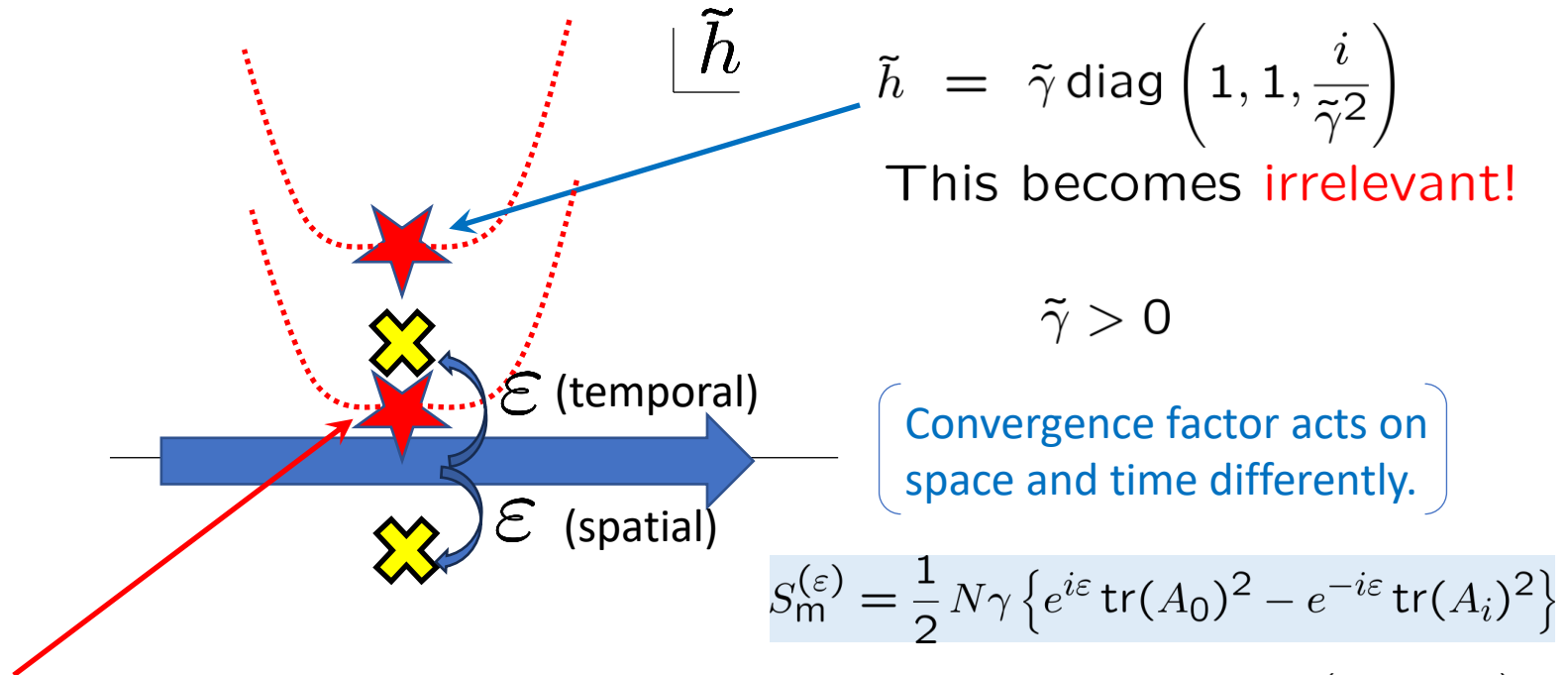
new saddle point appears near  $\tilde{h}_{\text{sing}} = \frac{\tilde{\gamma}}{2} \mathbf{1}$

$$Z_{\text{Pauli}} \sim \varepsilon^{-\frac{3}{2}D} (\times \varepsilon^6) \sim \varepsilon^{-\left(\frac{3}{2}D - 6\right)}$$

diverges as  $\varepsilon \rightarrow 0$

The new (relevant) saddle point appears, and the partition function diverges!

# Situation with the squashed Pauli



**new saddle point** appears near  $\tilde{h}_{\text{sing}} = c \mathbf{1} + (2c - \tilde{\gamma}) \text{diag} \left( 1, 1, \frac{i}{\tilde{\gamma}^2} \right)$

$$c = \frac{\tilde{\gamma}}{12} \left( 7 + \sqrt{1 - \frac{8i}{\tilde{\gamma}}} \right)$$

$$Z_{\text{s-Pauli}} \sim \epsilon^{-D} (\times \epsilon^3) \sim \epsilon^{-(D-3)} \quad \text{diverges as } \epsilon \rightarrow 0$$

The new (relevant) saddle point appears, and the partition function diverges!

# Physical meaning of the divergence

$$Z \sim \varepsilon^{-p}$$

Note: This does not mean that the model is ill defined.  
E.g., the expectation value  $\langle \text{tr}(A_\mu A^\mu) \rangle$  is finite.

Pauli

$$p \sim \frac{3}{2}D - 6$$



Partition function diverges  
faster for Pauli for  $D \gtrsim 6$

squashed Pauli

$$p \sim D - 3$$

Pauli has **3 nonvanishing internal d.o.f.**,  
while squashed Pauli has **only 2.**

This implies that Pauli thimble dominates  
in the cutoff model at  $\gamma > 0$  for  $D \gtrsim 6$ .

The diverging observables :

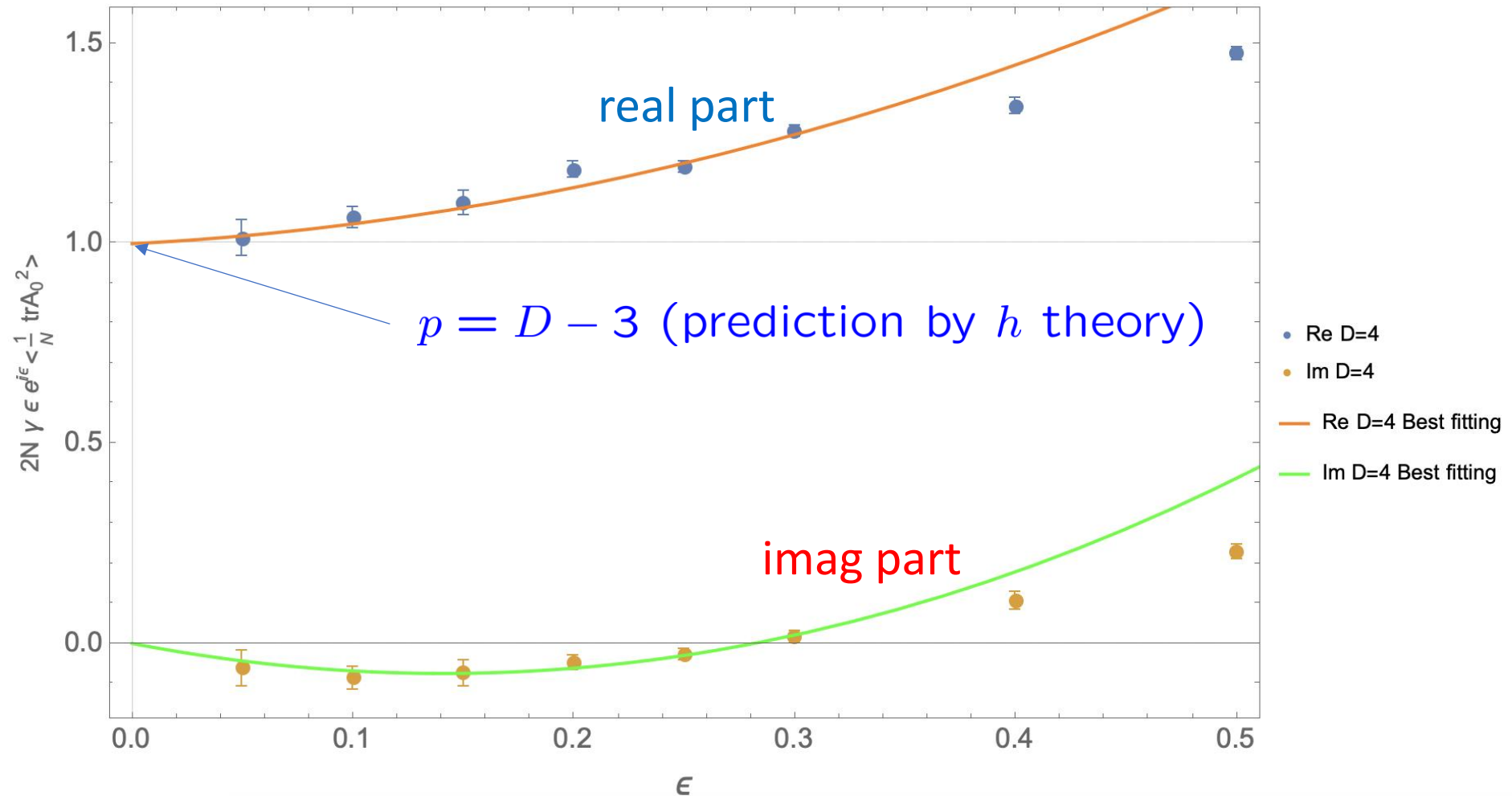
$$\left\langle \frac{1}{N} \text{tr}(A_0)^2 \right\rangle \sim -\frac{1}{2N\gamma e^{i\varepsilon}} \frac{\partial}{\partial \varepsilon} \log Z \sim \frac{p}{2N\gamma e^{i\varepsilon} \varepsilon}$$

Boosted configurations dominate the partition function.  
The cutoff artifact may well remain in the  $\varepsilon \rightarrow 0$  limit.

# Diverging

$$\left\langle \frac{1}{N} \text{tr} (A_0)^2 \right\rangle \sim \frac{p}{2N\gamma e^{i\varepsilon} \varepsilon}$$

$$\left\langle \frac{1}{N} \text{tr} (A_0)^2 \right\rangle \times 2N\gamma e^{i\varepsilon} \varepsilon \quad D = 4, \quad \text{squashed Pauli} \quad \gamma = 1.5$$



Consistent with  $Z \sim \varepsilon^{-(D-3)}$

# Classicalization for Pauli solution

- The new saddle point approaches  $\tilde{h}_{\text{sing}} = \frac{\tilde{\gamma}}{2} \mathbf{1}$

$$\frac{1}{\sqrt{D}} \frac{1}{N} \langle \text{tr} A_{\mu} A^{\mu} \rangle = \frac{1}{4} \langle \text{tr} \tilde{h} \rangle \sim \frac{3}{8} \tilde{\gamma}$$

- Fluctuations around the saddle point are suppressed at large D.

$$\lim_{D \rightarrow \infty} \frac{1}{\sqrt{D}} \frac{1}{N} \langle \text{tr} A_{\mu} A^{\mu} \rangle = \frac{3}{8} \tilde{\gamma} \quad (\text{classical result})$$

$$A_{\mu} = \begin{cases} \sqrt{\frac{\tilde{\gamma}}{8}} \sigma_{\mu} & \mu = 1, 2, 3 \\ 0 & \text{otherwise} \end{cases}$$

Classicalization for Pauli occurs at  $D = \infty$ .

# Hessian analysis around boosted Pauli

$$A_\mu = \begin{cases} \sqrt{\frac{\gamma}{8}} \sigma_\mu & \mu = 1, 2, 3 \\ 0 & \text{otherwise} \end{cases} \quad \xrightarrow{\text{boost in 1-direction}} \quad \begin{cases} A_0 = \sqrt{\frac{\gamma}{8}} \sinh \sigma \sigma_1 \\ A_1 = \sqrt{\frac{\gamma}{8}} \cosh \sigma \sigma_1 \\ A_2 = \sqrt{\frac{\gamma}{8}} \sigma_2 \\ A_3 = \sqrt{\frac{\gamma}{8}} \sigma_3 \\ A_i = 0 \quad \text{for } i \geq 4 \end{cases}$$

$$\delta S = \delta A_\mu^a H_{\mu\nu}^{ab}(A) \delta A_\nu^b \quad H_{\mu\nu}^{ab} = \frac{\partial^2 S(A)}{\partial A_\mu^a \partial A_\nu^b} \quad (\text{Hessian})$$

cutoff surface:  $\text{tr}(A_0)^2 + \text{tr}(A_i)^2 = \Lambda$

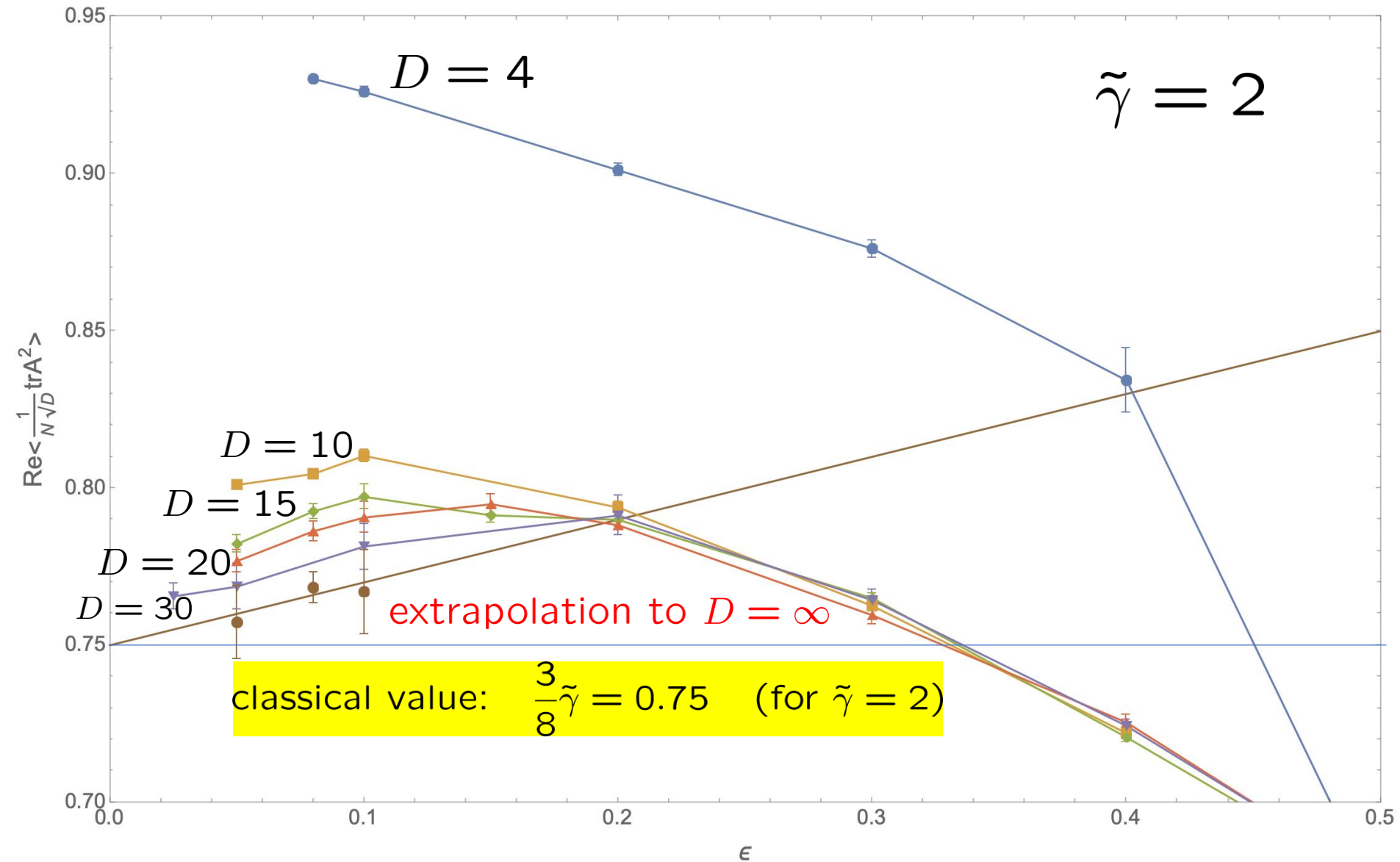
allowed fluctuations:  $\text{tr}(A_0 \delta A_0) + \text{tr}(A_i \delta A_i) = 0$

Eigenvalues of  $H$  in the  $(3D - 1)$ -dimensional subspace

- 2 finite (contributes to quantum corrections)
- 4 divergent (suppressed in the  $\varepsilon \rightarrow 0$  limit)
- $(3D - 7)$  zeroes (corresponding to broken symmetries)

**Classicalization occurs in the large  $D$  limit.**

# Classicalization for Pauli at $D = \infty$





# Hessian analysis around boosted squashed Pauli

$$A_\mu = \begin{cases} \sqrt{\frac{\gamma}{4}} \sigma_\mu & \mu = 1, 2 \\ 0 & \text{otherwise} \end{cases} \quad \xrightarrow{\text{boost in 1-direction}} \quad \begin{cases} A_0 = \sqrt{\frac{\gamma}{4}} \sinh \sigma \sigma_1 \\ A_1 = \sqrt{\frac{\gamma}{4}} \cosh \sigma \sigma_1 \\ A_2 = \sqrt{\frac{\gamma}{4}} \sigma_2 \\ A_i = 0 \quad \text{for } i \geq 3 \end{cases}$$

$$\delta S = \delta A_\mu^a H_{\mu\nu}^{ab}(A) \delta A_\nu^b \quad H_{\mu\nu}^{ab} = \frac{\partial^2 S(A)}{\partial A_\mu^a \partial A_\nu^b} \quad (\text{Hessian})$$

cutoff surface:  $\text{tr}(A_0)^2 + \text{tr}(A_i)^2 = \Lambda$

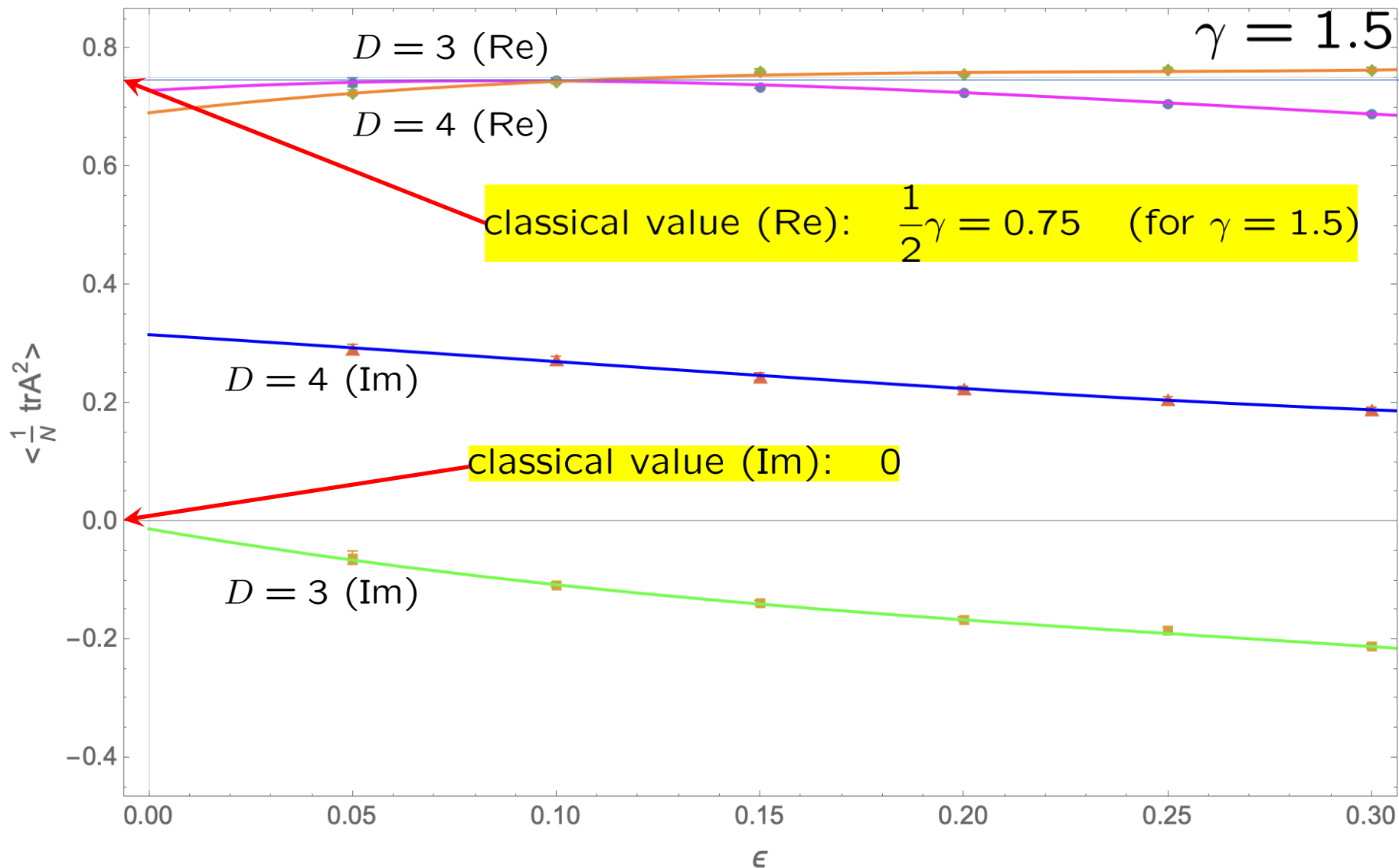
allowed fluctuations:  $\text{tr}(A_0 \delta A_0) + \text{tr}(A_i \delta A_i) = 0$

Eigenvalues of  $H$  in the  $(3D - 1)$ -dimensional subspace

- $(D - 3)$  finite (contributes to quantum corrections)
- 4 divergent (suppressed in the  $\varepsilon \rightarrow 0$  limit)
- $(2D - 2)$  zeroes (corresponding to broken symmetries)

Classicalization for squashed pauli occurs only at  $D = 3$ .

# Classicalization for squashed Pauli at $D = 3$



Classicalization occurs only at  $D = 3$ .

## 4. MC studies of the “gauge-fixed” model

Chou, JN, Tripathi, in preparation

# Saddle points in the gauge-fixed model

$$Z = \int dA e^{i(S_b + S_m)} \Delta_{\text{FP}}[A] \prod_{j=1}^d \delta(\text{tr}(A_0 A_j))$$

$$\Delta_{\text{FP}}[A] = \det \Omega$$

$$\Omega_{ij} = \text{tr}(A_0)^2 \delta_{ij} + \text{tr}(A_i A_j)$$

This represents the gauge fixing condition :  $\text{tr}(A_0 A_j) = 0$  for all  $j$

saddle point equation :

$$[A_\nu, [A^\nu, A_\mu]] = \gamma A_\mu + \frac{i}{N} \eta_{\mu\nu} \text{Tr} \left( \Omega^{-1} \frac{\partial \Omega}{\partial A_\nu} \right)$$

Using the  $\text{SO}(d)$  symmetry, we can impose :  $\text{tr}(A_i A_j) = 0$  for  $i \neq j$

$$[A_\nu, [A^\nu, A_\mu]] = (\gamma + i\kappa_\mu) A_\mu,$$

The effect of gauge-fixing appears here.

$$\kappa_i = \frac{2}{N \{ \text{tr}(A_0)^2 + \text{tr}(A_i)^2 \}}, \quad \kappa_0 = \sum_{i=1}^d \kappa_i.$$

$\kappa_\mu$  has to be determined in a self-consistent manner.

The FP determinant induces a mass-like term in the saddle point eq.

The  $\gamma \rightarrow 0$  limit may be smooth!

# Ansatz for the saddle points

$$[A_\nu, [A^\nu, A_\mu]] = (\gamma + i\kappa_\mu) A_\mu, \quad \kappa_i = \frac{2}{N\{\text{tr}(A_0)^2 + \text{tr}(A_i)^2\}}, \quad \kappa_0 = \sum_{i=1}^d \kappa_i.$$

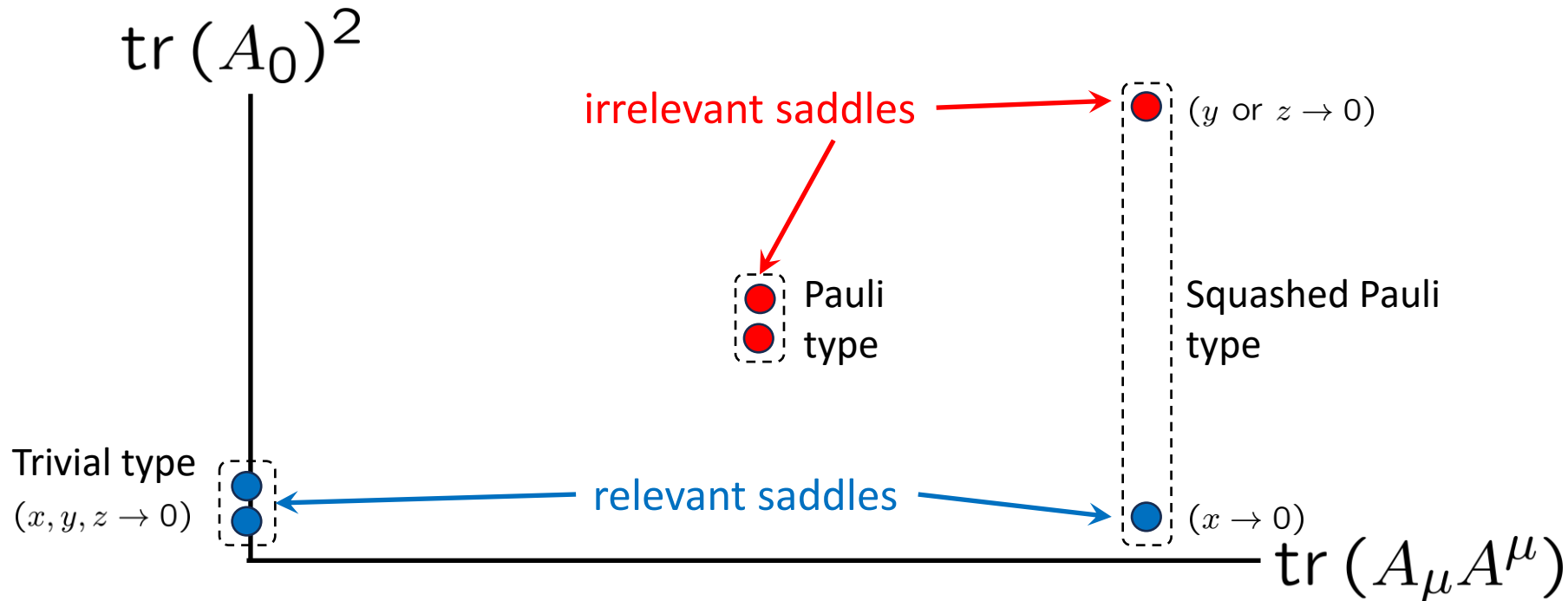
A natural ansatz:  
 $d \geq 4$

$$\begin{aligned} A_0 &= x \sigma_1, \\ A_1 &= y \sigma_2, \\ A_2 &= z \sigma_3, \\ A_j &= 0 \quad \text{for } 3 \leq j \leq d \end{aligned}$$

$A_0 \neq 0$  is required for finite  $\gamma$   
 since otherwise  $\Delta_{\text{FP}}[A] = 0$ .

solutions at  $\gamma > 0$

("type" =  $\gamma \rightarrow \infty$  behavior)



# The behavior of the solutions at $\gamma \rightarrow \infty$

For  $\gamma \rightarrow \infty$ , the solutions reduce to those of the gauge-unfixed model

$\gamma > 0$	$A_\mu = 0$	$A_\mu = \begin{cases} \sqrt{\frac{\gamma}{8}} \sigma_\mu & \mu = 1, 2, 3 \\ 0 & \text{otherwise} \end{cases}$	$A_\mu = \begin{cases} \sqrt{\frac{\gamma}{4}} \sigma_\mu & \mu = 1, 2 \\ 0 & \text{otherwise} \end{cases}$
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(trivial solution)

(Pauli solution)

(squashed Pauli solution)

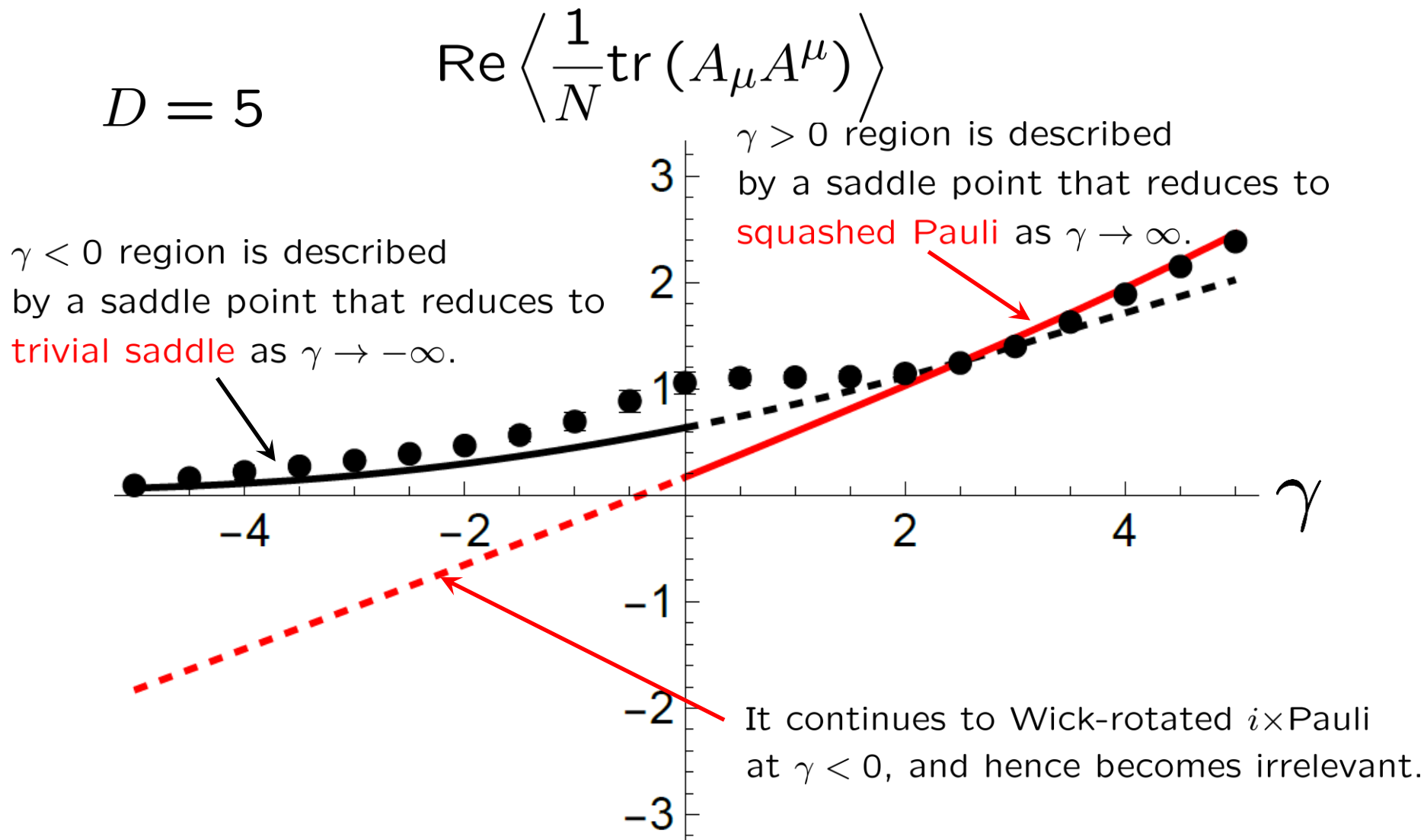
Recall, however, that solutions that are obtained by Wick rotation from above are **irrelevant** from the viewpoint of the Picard-Lefschetz theory.

$$\left\{ \begin{array}{l} A_0 = i \sqrt{\frac{\gamma}{8}} \sigma_1 \\ A_1 = \sqrt{\frac{\gamma}{8}} \sigma_2 \\ A_2 = \sqrt{\frac{\gamma}{8}} \sigma_3 \\ A_i = 0 \quad \text{for } i \geq 3 \end{array} \right. \quad \left\{ \begin{array}{l} A_0 = i \sqrt{\frac{\gamma}{4}} \sigma_1 \\ A_1 = \sqrt{\frac{\gamma}{4}} \sigma_2 \\ A_i = 0 \quad \text{for } i \geq 2 \end{array} \right.$$

Thus at large  $\gamma$ , relevant saddles should have  $A_0 \rightarrow 0$ .

**Pauli-type solution cannot be relevant.**

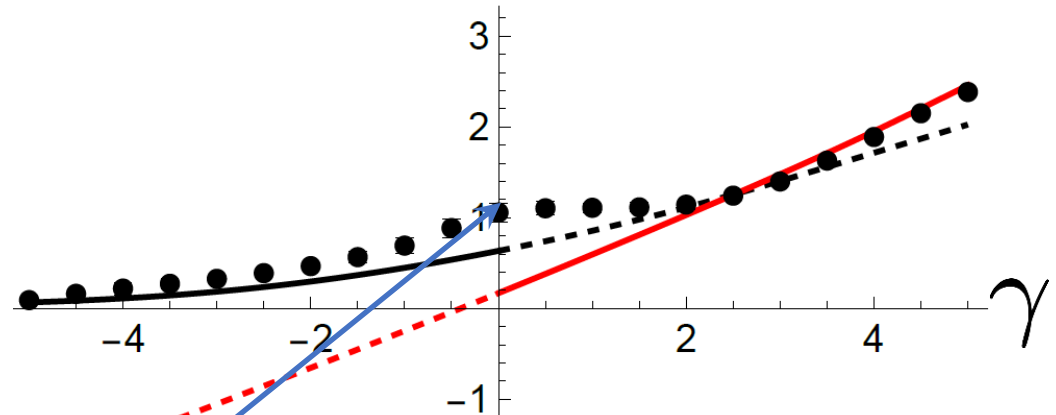
# Simulation results for the gauge fixed model (by the generalized Lefschetz thimble method)



Thus, the dominant saddle point for  $\gamma > 0$  is different from the cutoff model !

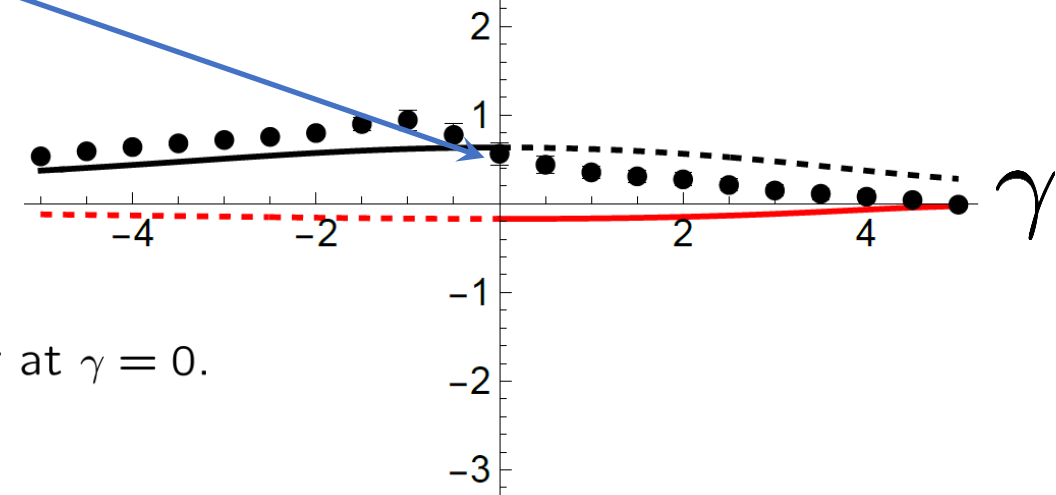
# Smooth behavior observed at $\gamma \sim 0$

$$D = 5 \quad \text{Re} \left\langle \frac{1}{N} \text{tr} (A_\mu A^\mu) \right\rangle$$



$\gamma \rightarrow 0$  limit seems to be smooth at least in this model

$$D = 5 \quad \text{Im} \left\langle \frac{1}{N} \text{tr} (A_\mu A^\mu) \right\rangle$$

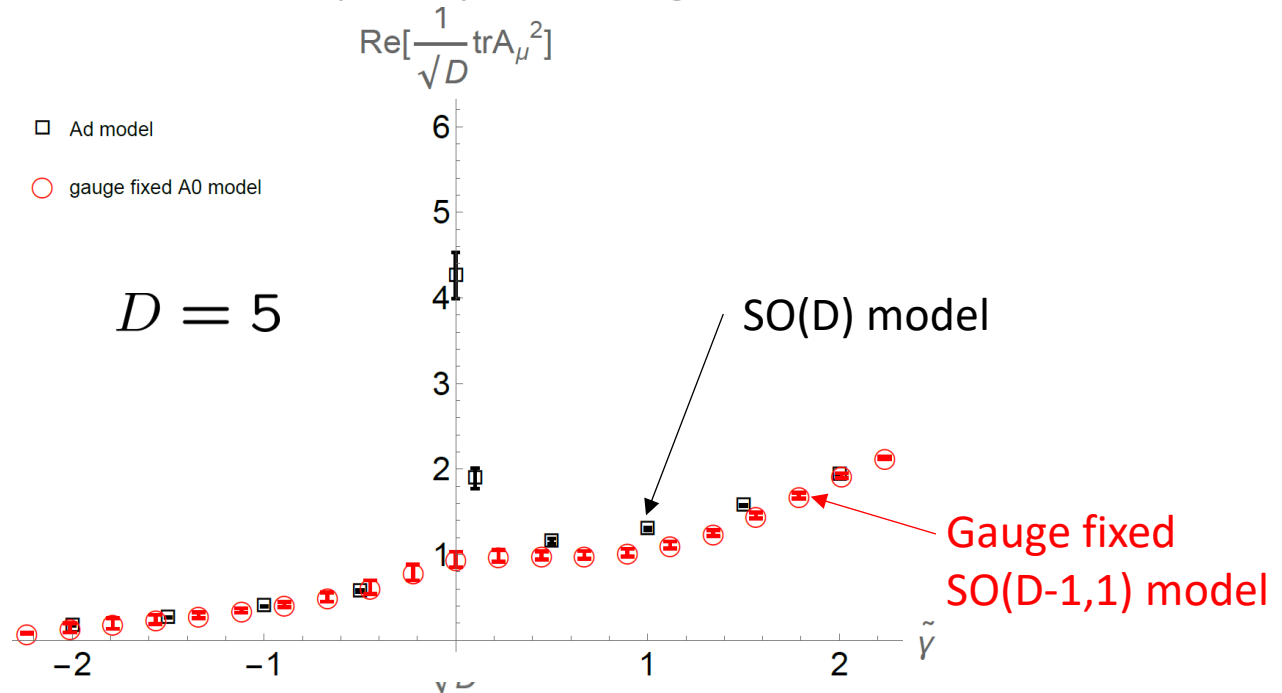


But, this might not be the case with SUSY and/or at large  $N$ .

Stokes phenomenon seems to occur at  $\gamma = 0$ .  
(Relevant saddle point changes.)

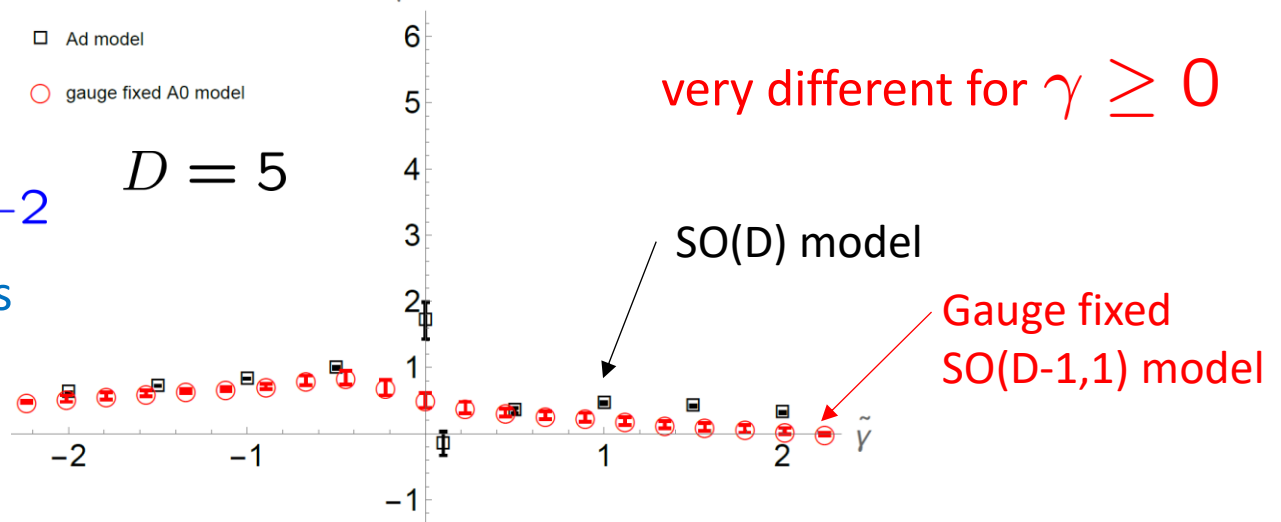


# Comparison with the SO(D) symmetric model obtained by replacing $A_0 = iA_D$

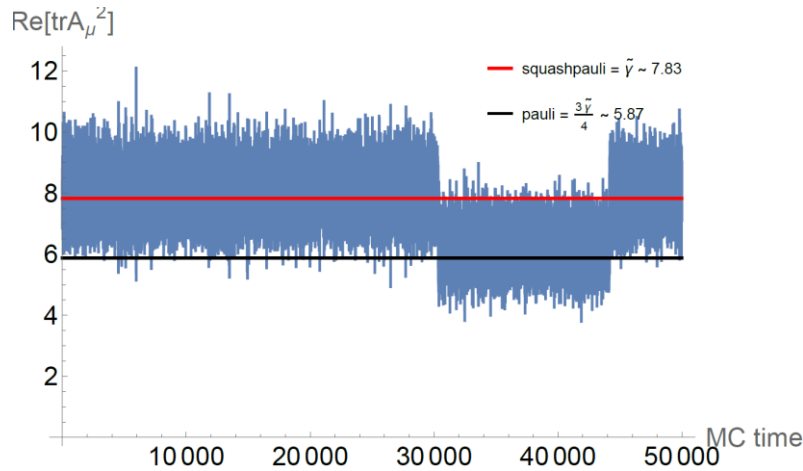


very different for  $\gamma \geq 0$

equivalence for  $\gamma \lesssim -2$   
trivial saddle dominates  
in this region



# Oscillating behavior in the SO(D) model at larger $\tilde{\gamma}$



Perturbative calculations around Pauli and squashed Pauli yield:

$$Z_{\text{Pauli}} \simeq \frac{\pi^{\frac{3(D+1)}{2}} \gamma^{\frac{3D}{2}-6} e^{-\frac{3i}{8}\gamma^2}}{2^{3(D-4)} \Gamma\left(\frac{D}{2}\right) \Gamma\left(\frac{D-1}{2}\right) \Gamma\left(\frac{D-2}{2}\right)}$$

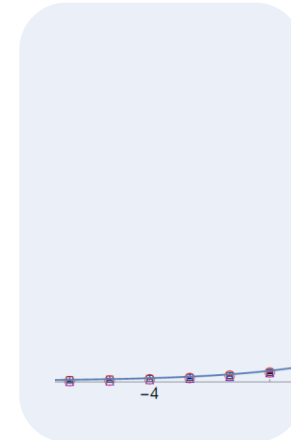
$$Z_{\text{s-Pauli}} \simeq \frac{\pi^{\frac{3D+2}{2}} \gamma^{\frac{D}{2}-1} e^{-\frac{i}{2}\gamma^2}}{2^{D-\frac{7}{2}} (-i)^{\frac{D-1}{2}} \Gamma\left(\frac{D}{2}\right) \Gamma\left(\frac{D-1}{2}\right)}$$

Due to the relative phase, interference occurs between P and sP.

At  $D = \infty$ , Pauli dominates over s-Pauli.

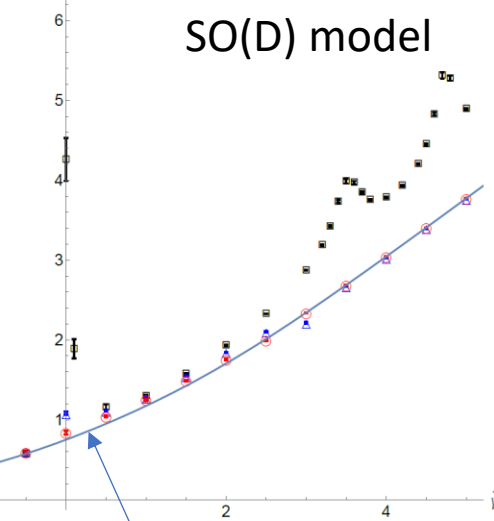
$D = 5, 10, 20$

□ D=5  
△ D=10  
○ D=20  
— h theory



$\text{Re}[-\frac{1}{\sqrt{D}} \text{tr} A_\mu^2]$

SO(D) model

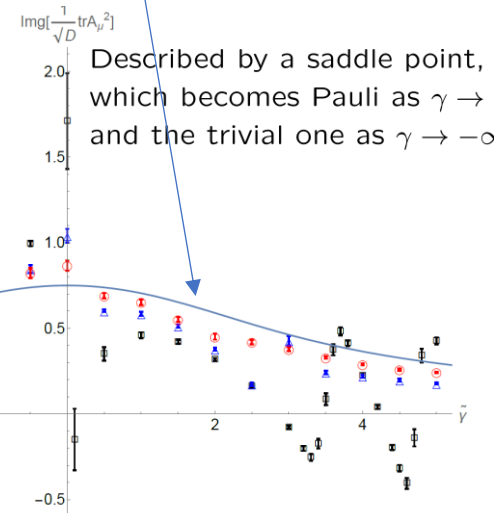


Prediction at  $D = \infty$



$\text{Img}[-\frac{1}{\sqrt{D}} \text{tr} A_\mu^2]$

Described by a saddle point, which becomes Pauli as  $\gamma \rightarrow \infty$  and the trivial one as  $\gamma \rightarrow -\infty$ .



## 5. Summary and discussions

# Summary

- **The type IIB matrix model** has diverging partition function due to Lorentz symmetry (represented by a noncompact group).
- In the cutoff model, **the Pauli solution** has more divergent partition function than the squashed Pauli, and hence **dominates**.  

Pauli has <b>3 nonvanishing internal d.o.f.</b> , while squashed Pauli has only <b>2</b> .
---
- **The cutoff model** suffers from a severe artifact due to Lorentz symmetry breaking. (Classicalization at  $D=\infty$  Pauli,  $D=3$  sPauli.)
- We have proposed a new definition of type IIB matrix model without Lorentz symmetry breaking using the gauge fixing. In the gauge-fixed model, Pauli solution cannot appear at large  $\gamma$ .

Gauge fixing is crucial in determining the dominant saddle point.

# Future prospects

- What happens in the SUSY case and/or at larger N.

Does (3+1)-dimensional expanding space-time emerge in the 1)  $N \rightarrow \infty$ , 2)  $\gamma \rightarrow 0$  limit ?

- SUSY case

1/D expansion cannot be applied (SUSY cannot be respected), but **numerical simulation is doable**. N=2 case is on-going.

- larger N

The computational cost of the **generalized Lefschetz thimble method** grows with N as  $O(N^6)$ . But we may still do N=4,8,16,...

- SUSY and large N

The Pfaffian seems to **prefer collapsed configurations**, but it becomes **zero** for configurations with **not more than 2 extended directions**.

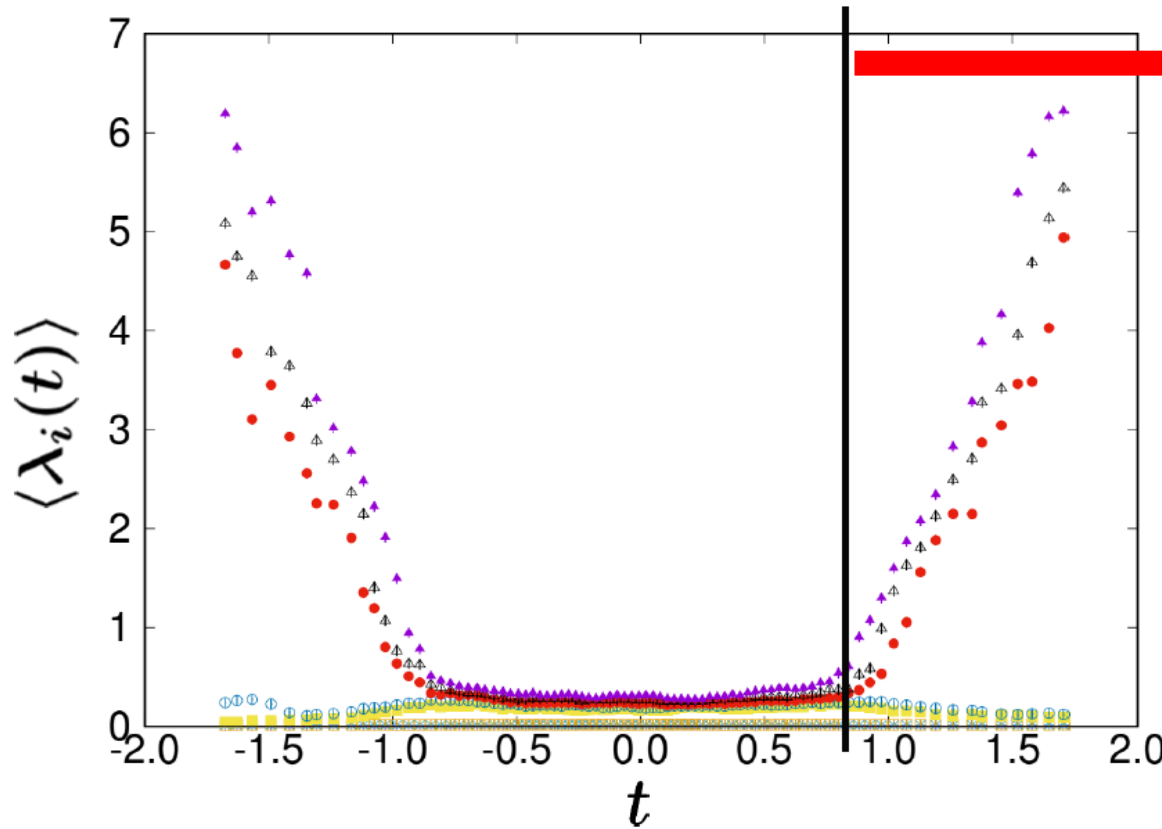
➡ 3d space ? **complex Langevin method (less flexible but cheaper)**

# Recent results from complex Langevin simulation

(gauge-unfixed model)

Anagnostopoulos, Azuma, Hatakeyama, Hirasawa, JN, Papadoudis, Tsuchiya, in preparation

The extent in 9 directions v.s. time



Emergence of expanding (3+1)D space-time

$$N = 96$$

$$\gamma = 4$$

$$m_f = 3.5$$

$$d = 5, \quad \xi = 16$$

5 directions are suppressed by hand for technical reasons.

The 4th direction becomes small at late times spontaneously !